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Solar Energy xxx (2007) xxx-xxx



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² Modeling of the optimum tilt of a solar chimney for maximum air flow

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Received 18 July 2006; received in revised form 5 February 2007; accepted 6 March 2007

Communicated by: Associate Editor S.A. Sherif

9 Abstract

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The aim of this work is to develop a mathematical model to determine the tilt that maximizes natural air flow inside a solar chimney using daily solar irradiance data on a horizontal plane at a site. The model starts by calculating the hourly solar irradiation components (direct, diffuse, ground-reflected) absorbed by the solar chimney of varying tilt and height for a given time (day of the year, hour) and place (latitude). In doing so it computes the transmittance and absorbance of the glazing for the various solar irradiation components and for various tilts. The model predicts the temperature and velocity of the air inside the chimney as well as the temperatures of the glazing and the black painted absorber. Comparisons of the model predictions with CFD calculations delineate the usefulness of the model. In addition, there is a good agreement between theoretical predictions and experiments performed with a 1 m long solar chimney

17 at different tilt positions.

18 © 2007 Published by Elsevier Ltd.

19 *Keywords:* Solar chimney; Chimney effect; Natural ventilation; Solar air heater; Tilt; Maximum flow 20

21 1. Introduction

22 Solar chimneys differ from conventional chimneys in 23 that their southern wall (for the north hemisphere) is 24 replaced by a transparent sheet, i.e. glazing, that allows 25 the collection and use of solar irradiation. Many works, 26 especially the last two decades, have illustrated the advantages in using solar chimneys accounting also for their low 27 maintenance cost and superb durability. Solar chimneys 28 29 have been traditionally used in agriculture for air renewal 30 in barns, silos, greenhouses, etc. as well as in drying of 31 crops, grains, fruits or wood (e.g. Garg, 1987; Das and 32 Kumar, 1989; Ekechukwu and Norton, 1995; Vlachos 33 et al., 2002). Another popular application is for natural 34 ventilation in buildings in order to improve the quality of 35 indoors air and increase the comfort index for inhabitants

0038-092X/\$ - see front matter @ 2007 Published by Elsevier Ltd. doi:10.1016/j.solener.2007.03.001

(e.g. Kumar et al., 1998; Ziskind et al., 2002; Ding et al., 36 2004; Bansal et al., 2005). Having in mind climatization 37 and energy conservation in buildings, efforts have also been 38 made to evaluate the performance of special chimney con-39 figurations, such as solar roof collectors and Trombe-walls 40 (e.g. Gan, 1998; Sànchez et al., 2003; Ong and Chow, 2003; 41 Khedari et al., 2000, 2003; Heras et al., 2005) as well as 42 other hybrid constructions involving inclined, vertical or 43 horizontal heated walls with cooling cavities (e.g. Raman 44 et al., 2001; Jiang and Chen, 2003; Kazansky et al., 2003; 45 Dai et al., 2003). 46

Most published works deal with solar chimneys fixed at 47 a specific inclination, usually vertical, as these are easier to 48 construct and operate. To overcome the diurnal variability 49 of solar irradiance, some of them used heating elements to 50 maintain either uniform wall temperature or wall heat flux 51 and examined only the heat transfer and fluid mechanics 52 performance of the chimney (e.g. Bouchair and Fitzgerald, 53 1988; Bouchair, 1994; Moshfegh and Sandberg, 1999; 54 Afonso and Oliveira, 2000; Chen et al., 2003). These studies -55

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Nomenclature

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Latin s	ymbols	U	overall heat transfer coefficient
a A	absorptance	W	width of the chimney gap
A C	discharge coefficient	Graak	symbols
c_d	specific heat	α and β coefficients in Eq. (9)	
d	depth of the chimney gap	δ and μ	declination (angular position of the sun at solar
и Д.,	hydraulic diameter of the chimney	0	noon)
\mathcal{D}_{H} f	wall friction coefficient	ę	emmitance of the black wall
G	solar constant (1367 W/m^2)	n, and	n_{2} refraction indexes of air and glass respectively
h	convective heat transfer coefficient	θ	angle
hour	hour of the day	$\hat{\theta}_1$	angle of irradiation incidence
H	daily irradiation on a horizontal plane	θ_2	angle of refraction
Hah	height difference between outlet and inlet of the	λ	thermal conductivity of air
ch	chimnev	u	viscosity
H_0	daily extraterrestrial irradiation on a horizontal	ρ	density
0	plane	σ	Stefan-Boltzmann constant (= 5.6697×10^{-8} W/
Ι	hourly irradiation on a horizontal plane		$m^2 K^4$)
Κ	extinction coefficient of the glass	τ	glass transmittance,
$k_{\rm in}$ and	k_{out} inlet and outlet pressure loss coefficients	$ au_{ m r}$	average of r_{\perp} and r_{\parallel}
k_{T}	clearness index	υ	average air velocity inside the chimney
ℓ	path length of irradiation through the glass	ϕ	latitude of the site (angular distance from the
L	length of the chimney		equator)
п	day of the year (1 to 365)	ω	hour angle
Nu	Nusselt number		
r	ratio of the hourly irradiation over the daily	Subscripts	
	irradiation	α	absorber
rg	diffuse reflectance of the surroundings	αbs	absorption
r_{\perp}	perpendicular component of unpolarized irradi-	air	air
	ation	bw	black wall
r_{\parallel}	parallel component of unpolarized irradiation	dif	diffuse
$R_{\rm b}$	ratio of the direct irradiation on a tilted plane	dir	direct
	over that on the horizontal plane	g	glazing
Ra	Rayleigh number	gap	chimney gap
Ra_{c}	critical Rayleigh number	0	ambient conditions
Re	apparent Reynolds number	ref	reflection
S	slope of the chimney with respect to the horizon-	s T	sunset
T	tal plane	Т	tilted plane
T	temperature		

showed that there are distinctly different flow patternsbetween narrow and wide chimney gaps and that the ratioof chimney length/gap influences the air flow rate.

59 Awbi and Gan (1992) obtained analytically the air tem-60 perature and flow rate profiles along a Trombe wall, con-61 sidering a uniform wall temperature. The same authors 62 employed also CFD codes to simulate the air flow and heat transfer in a chimney of varying gap width (for large gaps 63 64 3D simulations were indispensable). Both analytical and numerical results were in good agreement with earlier data. 65 66 Bansal et al. (1993) developed a steady state analytical 67 model for uniform wall temperature applied to a solar system consisting of a solar air heater connected to a conven-68 69 tional chimney. Andersen (1995) derived a set of equations to predict the natural ventilation in a room with small 70 71 openings based on the pressure model. Gan (1998) and Gan and Riffat (1998) used 3D CFD techniques to study 72 the parameters that influence the performance of a Trombe 73 wall. An interesting result of these studies was that ventila-74 tion rates increased along with the thickness of the interior 75 wall. Moshfegh and Sandberg (1999), investigated air 76 movement behind photovoltaic panels using a 2D CFD 77 code coupled with a standard $k-\varepsilon$ turbulence model and a 78 79 wall function. Their predictions of air velocity and temperature distributions were in accord with their experimental 80 results. A similar 2D CFD approach was also adopted by 81 Rodrigues et al. (2000) who provided detailed calculations 82 of the velocity and temperature profiles in the chimney. 83

84 Using the concept of a thermal resistance network, Ong 85 and Chow (2003) developed an analytical model to exam-86 ine the effects of air gap and solar irradiation intensity on 87 the performance of different chimneys assuming uniform 88 heat flux on the heated wall. Many of the above studies 89 provided evidence that for chimneys with gap-to-length 90 ratio less than or close to 1:10, the temperature can be 91 assumed uniform across the chimney gap and so 2D mod-92 els can give reasonably accurate predictions.

93 Solar chimneys employing inclined collectors can evi-94 dently exploit more the incident irradiation to enhance 95 air flow in the chimney. As the inclination of the chimney 96 varies, two things occur that work in opposite directions 97 with respect to the air flow rate. A higher inclination results 98 in a higher exposure of the wall to solar irradiation and 99 hence yield higher heat utilization and more intense buoy-100 ant airflow. On the other hand, tilting the chimney reduces 101 the effective pressure head of the chimney and so dimin-102 ishes air flow. It is apparent that there must be an optimum 103 tilt that leads to the highest flow rate, compromising these 104 two effects. Although there are e few studies coping with 105 the effect of inclination on a chimney performance, they 106 usually involve heating means other than solar irradiance 107 to achieve uniform wall heat flux (e.g. Moshfegh and Sand-108 berg, 1999; Chen et al., 2003) and so a parametric analysis 109 with respect to the temporal variability of solar irradiation 110 is not possible.

111 To our knowledge, there are only two earlier studies that 112 examined systematically the effect of inclination for chim-113 neys where the absorbed heat flux depends on the diurnal 114 and seasonal variations of solar irradiation. The first is 115 the work by Prasad and Chandra (1990) who performed 116 numerical calculations and also did experiments for a solar 117 chimney 1.5 m long and with 20 mm gap width. Their 118 model, though detailed for the momentum and heat trans-119 fer in the fluid, had certain drawbacks: did not account for 120 heat losses, required knowledge of the ratio of diffuse/total 121 irradiation and, finally, assumed that the transmittance of 122 the glazing and absorptance of the black wall were unity 123 although it is known that these quantities vary with inclina-124 tion (Duffie and Beckman, 1991). The agreement between 125 predictions and experiments was rather poor but their find-126 ing that the optimum tilt for maximum irradiance uptake 127 (i.e. maximum air temperature), is distinctly different than 128 the optimum tilt for maximum air velocity in the chimney 129 was significant. They calculated optimum tilt angles for 130 maximum air velocity to oscillate periodically throughout 131 the year between a low value, 52°, in summer months 132 and a high value, 72°, in winter months (for Calcutta, 133 India). A much simpler treatment of solar irradiation and 134 glazing optical properties was employed by Hamdy and 135 Fikry (1998) for Alexandria (Egypt) and summer months. 136 For these particular conditions, an optimum tilt around 137 60° was estimated for maximum air flow.

Data of solar irradiation at a site that are easily accessi-ble by a design engineer usually refer to monthly averagedaily values of total irradiation on a horizontal plane,

e.g. ELOT (1991). Therefore for designing a solar chimney, 141 horizontal irradiation data have first to be transformed to 142 irradiation data at a slope. For accurate estimations, it is 143 important to base the design on hourly values of solar irra-144 diation which must then be determined from the available 145 daily values. In any case, it is necessary to decompose the 146 total irradiation arriving at the sloped surface into its 147 major components (direct, diffuse and ground-reflected) 148 since for each of them the optical properties (transmittance 149 and absorptance) of the glass cover varies differently with 150 the tilt. As far as we know, there is no prior work that deals 151 with the estimation of the optimum tilt of a solar chimney 152 for maximizing air flow starting from data of daily total 153 solar irradiation on an horizontal plane and taking into 154 account all the above considerations. It is indeed the scope 155 of this study to propose an engineering model that can 156 cope with this task. The term engineering denotes a simpli-157 fied model adequate for design purposes and field applica-158 tions which does not employ detailed 2D/3D fluid 159 mechanics and heat transfer calculations. 160

In the following, the setup of the engineering and CFD 161 models is presented first. Next, the solar chimney construction and operation are outlined. Finally, theoretical predictions from the models are compared and discussed against each other and against experimental results. 165

2. Theory

2.1. Engineering model 167

In the analysis below, it is assumed that the incident 168 solar irradiation is sufficient to bring the chimney's body 169 to its steady state temperature. The input data to the model 170 are divided in five categories: (a) chronological (day of the 171 year, hour) and geographical (latitude), (b) meteorological 172 (monthly average daily total irradiation on an horizontal 173 174 plane, monthly average daily ambient temperature), (c) geometrical (dimensions of chimney, thicknesses of glazing 175 and insulation material), (d) optical/irradiation properties 176 of the construction materials (refractive index and extinc-177 tion coefficient of the glazing, absorptance and emmitance 178 of the black surfaces), and (e) physical properties of air and 179 insulation materials for calculating heat losses). The chim-180 ney tilt and length are treated as variables in the range 181 30-90° (angles from the horizontal plane) and 1-12 m, 182 respectively. Physical properties of air are taken from 183 VDI-Wärmeatlas (1991). Data for monthly average daily 184 total irradiation and monthly average ambient temperature 185 are taken from ELOT (1991) – the Greek Organization of 186 Standardization - for Serres, a city in North Greece where 187 188 also the experimental tests are performed. The ELOT data agree reasonably well with measurements from the meteo-189 190 rological station of TEI – Serres with only a $\sim 3\%$ annual deviation (Karapantsios et al., 1999). 191

The model consists of three basic subroutines. The first 192 one estimates the solar irradiation components (direct, 193 diffuse and ground-reflected) that hit the surface of the 194

Please cite this article in press as: Sakonidou, E.P. et al., Modeling of the optimum tilt of a solar chimney for maximum air flow, Sol. Energy (2007), doi:10.1016/j.solener.2007.03.001

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chimney on an hourly basis, at varying tilt and length. For
this calculation, only the chronological, geographical and
meteorological information mentioned above is needed.
The relations below, unless differently stated, are taken
from Duffie and Beckman (1991).

200 The total daily irradiation on a horizontal plane, H, is 201 customary expressed as the sum of two components: the 202 direct (beam) irradiation and the diffuse irradiation from 203 the sky

$$206 \quad H = H_{\rm dir} + H_{\rm dif} \tag{1}$$

The daily extraterrestrial solar irradiation H_0 on a horizontal plane is given as

$$H_{0} = \frac{24 \times 3600}{\pi} G_{\rm sc} \left(1 + 0.033 \cos \frac{360n}{365} \right)$$

210
$$\times \left[\cos \phi \cos \delta \sin \omega_{\rm s} + \frac{\pi \omega_{\rm s}}{180} \sin \phi \sin \delta \right]$$
(2)

211 where $G_{\rm sc}$ is the solar constant (1367 W/m²), *n* is the day of 212 the year (1–365), ϕ is the latitude of the site (angular dis-213 tance from the equator, for Serres ($\phi = 41$), δ is the decli-214 nation (angular position of the sun at solar noon) and $\omega_{\rm s}$ 215 is the sunset hour angle given as

217
$$\omega_{\rm s} = \arccos(-\tan\phi\tan\delta)$$
 (3)

218 The declination δ is found from the equation:

220
$$\delta = 23.45 \sin\left(360 \frac{284 + n}{365}\right)$$
 (4)

The daily extraterrestrial solar irradiation H_0 is related with the daily total irradiation H (input variable to the code), via the clearness index $k_{\rm T}$:

$$k_{\rm T} = \frac{H}{H_0} \tag{5}$$

226 Knowing the value of the clearness index, one can calcu-227 late the diffuse component, H_{dif} , as follows for $\omega_{\text{s}} \leq 81.4^{\circ}$:

$$\frac{H_{\rm dif}}{H} = \begin{cases} 1.0 - 0.2727k_{\rm T} + 2.4495k_{\rm T}^2 \\ -11.9514k_{\rm T}^3 + 9.3879k_{\rm T}^4 & \text{for } k_{\rm T} < 0.715 \\ 0.143 & \text{for } k_{\rm T} \ge 0.715 \end{cases}$$

229

232

230 whereas for $\omega_{\rm s} > 81.4^{\circ}$

$$\frac{H_{\rm dif}}{H} = \begin{cases} 1.0 + 0.2832k_{\rm T} - 2.5557k_{\rm T}^2 + 0.8448k_{\rm T}^3 \\ \text{for } k_{\rm T} < 0.722 \\ 0.143 \quad \text{for } k_{\rm T} \ge 0.722 \end{cases}$$
(7)

Then, the direct daily component can be computed from Eq. (1). The ratio of the total hourly irradiation, I, over the total daily irradiation, H, is given by

$$238 \quad r_t = \frac{I}{H} \tag{8}$$

and can be found from the relation

242
$$r_t = \frac{\pi}{24} (\alpha + \beta \cos \omega) \frac{\cos \omega - \cos \omega_s}{\sin \omega_s - \frac{\pi \omega_s}{180} \cos \omega_s}$$
(9)

where the coefficients α and β are given by 243

$$\alpha = 0.409 + 0.5016\sin(\omega_{\rm s} - 60) \tag{10}$$

$$\beta = 0.6609 - 0.4767\sin(\omega_{\rm s} - 60) \tag{11}$$

and the hour angle ω that appears above is given by 246

$$\omega = (hour - 12) * \frac{360}{24} \tag{12} 248$$

where *hour* denotes the hour of the day (input variable). 249 Likewise, the ratio of the hourly diffuse irradiation, I_{dif} , 250 over the daily diffuse irradiation, H_{dif} , is given as 251 252

$$r_{\rm dif} = \frac{I_{\rm dif}}{H_{\rm dif}} \tag{13}$$

where

$$r_{\rm dif} = \frac{\pi}{24} \frac{\cos \omega - \cos \omega_{\rm s}}{\sin \omega_{\rm s} - \frac{\pi \omega_{\rm s}}{180} \cos \omega_{\rm s}} \tag{14}$$

Then, I and I_{dif} are calculated from Eqs. (8) and (13), 258 respectively, whereas the hourly direct irradiation, I_{dir} , is 259 computed as 260

$$I_{\rm dir} = I - I_{\rm dif} \tag{15} \quad 262$$

A customary approach for irradiation estimations on 263 sloped surfaces is to consider an isotropic 2D model for 264 the diffuse irradiation (Liu and Jordan, 1963) and also 265 assume that the reflecting surfaces are diffuse and not spec-266 ular reflectors. Recent studies, (e.g. Badescu, 2002) have 267 shown that isotropic 3D models perform better than the 268 Liu-Jordan isotropic 2D model which seems to overesti-269 mate the diffuse and underestimate the ground reflected 270 solar irradiation component, respectively. For sites around 271 our latitude or smaller, the diffuse and reflected compo-272 nents are usually much less than the direct solar irradiation 273 component. In addition, for the most common chimneys 274 inclinations (close to vertical) the differences between 2D 275 and 3D models diminish and, therefore, can be safely 276 ignored. For these reasons the present calculations utilize 277 the well-known Liu–Jordan isotropic 2D model. In this 278 279 case, the total irradiation on a surface tilted at slope s, is given by Sukhatme (1984) 280281

$$I_{\rm T} = I_{\rm dir}R_{\rm b} + I_{\rm dif}\left(\frac{1+\cos s}{2}\right) + Ir_{\rm g}\left(\frac{1-\cos s}{2}\right) \tag{16}$$

On the RHS of Eq. (16), the first term represents the 284 direct component, the second term the diffuse component 285 and the third term the component that is reflected from 286 the surroundings. In (16), $r_{\rm g}$ is the diffuse reflectance of 287 the surroundings (usually around 0.25) and $R_{\rm b}$ is the ratio 288 of direct irradiation on the tilted surface over that on the 289 horizontal plane. R_b for the northern hemisphere is given 290 291 by Sukhatme (1984):

$$R_{\rm b} = \frac{\cos(\phi - s)\cos\delta\cos\omega + \sin(\phi - s)\sin\delta}{\cos\phi\cos\delta\cos\omega + \sin\phi\sin\delta} \tag{17}$$

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Thus, the average hourly irradiation components on a tilted surface that enter in the calculations of the solar chimney are

$$I_{\mathrm{T,dir}} = I_{\mathrm{dir}} \cdot R_{\mathrm{b}} \tag{18}$$

$$I_{\rm T,dif} = I_{\rm dif} \cdot \left(\frac{1+\cos s}{2}\right) \tag{19}$$

$$I_{\mathrm{T,ref}} = I \cdot 0.25 \cdot \left(\frac{1 - \cos s}{2}\right) \tag{20}$$

The second subroutine of the model evaluates the transmittance, τ , and absorptance, a_g , of the glazing for the various components of the incident solar irradiation. The transmittance for the direct irradiation component, τ_{dir} , is approximately given by the product:

$$306 \quad \tau_{\rm dir} \cong \tau_{\alpha \rm bs, dir} \cdot \tau_{\rm ref, dir} \tag{21}$$

307 where $\tau_{\alpha bs,dir}$ denotes the ratio of the transmitted versus the 308 incident irradiation where only absorption losses have been 309 considered and $\tau_{ref,dir}$ denotes the transmittance of initially 310 unpolarized irradiation where only reflection losses have 311 been considered.

312 In Eq. (21), $\tau_{\alpha bs, dir}$ is given as

$$\tau_{\alpha bs, dir} = \exp\left(-\frac{K\ell}{\cos\theta_2}\right) \tag{22}$$

where *K* is the extinction coefficient of the glass that varies 315 from approximately 4 m^{-1} for "water white" glass to 316 approximately 32 m^{-1} for poor (greenish cast of edge) glass 317 (Duffie and Beckman, 1991). In this work, $K = 10 \text{ m}^{-1}$. 318 Moreover, ℓ designates the path length of irradiation 319 through the glass which in effect is the thickness of the 320 glass. In this work, $\ell = 0.004$ m. Finally, θ_2 is the angle 321 of refraction, which is calculated from the expression: 322

$$\frac{\eta_1}{\eta_2} = \frac{\sin \theta_1}{\sin \theta_2} \tag{23}$$

where η_1 and η_2 are the refraction indexes of air and glass, 325 respectively; in this work, $\eta_1 = 1$ and $\eta_2 = 1.526$ (Duffie 326 and Beckman, 1991). Furthermore, θ_1 is the angle of incidence calculated as 328

$$\theta_1 = \arccos\left[\sin(\phi - s)\sin\delta + \cos(\phi - s)\cos\delta\cos\omega\right] \quad (24) \quad \widetilde{331}$$

The parameter τ_r is the average of two components 332

$$\tau_{\rm r,dir} = \frac{1}{2} \left(\frac{1 - r_{\parallel}}{1 + r_{\parallel}} + \frac{1 - r_{\perp}}{1 + r_{\perp}} \right)$$
(25) 334

where r_{\perp} represents the perpendicular component and r_{\parallel} 335 the parallel component of unpolarized irradiation given by 336





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$$r_{\perp} = \frac{\sin^2(\theta_2 - \theta_1)}{\sin^2(\theta_2 + \theta_1)}$$
(26)

$$r_{\parallel} = \frac{\tan^2(\theta_2 - \theta_1)}{\tan^2(\theta_2 + \theta_1)}$$
(27)

The diffusion component, τ_{dif} , and the reflection component, τ_{ref} , of the glass transmittance, are calculated in the same manner as τ_{dir} but with the slope *s* in Eq. (24) replaced by the diffusion angle, θ_{dif} , and the reflection angle, θ_{ref} , respectively, defined as

345
$$\theta_{\rm dif} = 59.7 - 0.1388 \cdot s + 0.001497 \cdot s^2$$
 (28)

346 and

348
$$\theta_{\rm ref} = 90 - 0.5788 \cdot s + 0.002693 \cdot s^2$$
 (29)

Accordingly, the components of the glass absorptance: direct, $a_{g,dir}$, diffuse, $a_{g,dif}$ and reflected, $a_{g,ref}$, are given by the following approximate relations:

$$a_{\rm g,dir} \simeq 1 - \tau_{\rm abs,dir} \tag{30}$$

 $a_{\text{g,dif}} \cong 1 - \tau_{\text{xbs,dif}} \tag{31}$

 $a_{g,ref} \simeq 1 - \tau_{\alpha bs,ref} \tag{32}$

For conciseness, the products $(\tau \cdot I)_T$ and $(a_g \cdot I)_T$ will henceforth denote the following quantities:

$$(\tau \cdot I)_{\rm T} = \tau_{\rm dir} \cdot I_{\rm T, dir} + \tau_{\rm dif} \cdot I_{\rm T, dif} + \tau_{\rm ref} \cdot I_{\rm T, ref}$$
(33)

357
$$(a_{g} \cdot I)_{T} = a_{g,dir} \cdot I_{T,dir} + a_{g,dif} \cdot I_{T,dif} + a_{g,ref} \cdot I_{T,ref}$$
 (34)

These are quantities that appear in the heat balance equations of the chimney and therefore need be evaluated first when running the code.

361 The third subroutine of the model solves the overall 362 energy balance of the chimney in the form of a system of three algebraic equations describing the heat exchange 363 across the black wall (absorber), the glazing and the air 364 inside the chimney, respectively. Fig. 1b shows a schematic 365 representation of the solar chimney configuration including 366 367 most critical elements. The heat exchange equations are 368 (temperatures in Kelvin)

$$a_{bw}(\tau \cdot I)_{T}A_{g}$$

$$= U_{bw} \cdot A_{bw} \cdot (T_{bw} - T_{0}) + h_{bw} \cdot A_{bw} \cdot (T_{bw} - T_{air})$$

$$+ \varepsilon \cdot \sigma \cdot A_{g} \cdot (T_{bw}^{4} - T_{g}^{4})$$

$$(a_{g} \cdot I)_{T}A_{g} + \varepsilon \cdot \sigma \cdot A_{g} \cdot (T_{b}^{4} - T_{g}^{4})$$

$$(35)$$

$$= h_{g} \cdot A_{g} \cdot (T_{g} - T_{air}) + U_{g} \cdot A_{g} \cdot (T_{g} - T_{0})$$
(36)

$$h_{\text{bw}} \cdot A_{\text{bw}} \cdot (T_{\text{g}} - T_{\text{air}}) + h_{\text{g}} \cdot A_{\text{g}} \cdot (T_{\text{g}} - T_{\text{air}})$$

$$= 2 \cdot c_{p,\text{air}} \cdot \rho_{\text{air}} \cdot A_{\text{gap}} \cdot \upsilon \cdot (T_{\text{air}} - T_{0})$$
(37)

371 In these equations, the three unknowns are T_{bw} the aver-372 age temperature of the black wall, T_{air} the average air temperature in the chimney and T_g the average glass 373 374 temperature. T_0 , is the ambient temperature which is an 375 input variable. Moreover, A_{bw} is the surface area of the 376 black wall, A_{g} is the surface area of the glass cover and A_{gap} 377 is the cross sectional area of the chimney gap. Further-378 more, $U_{\rm bw}$ is the overall heat transfer coefficient between 379 the black wall and the surroundings (in this work,

 $U_{\rm bw} = 0.9 \text{ W/m}^2 \text{ K}$ for a typical insulation thickness of 380 5 cm with thermal conductivity of 0.045 W/m K and mod-381 erate ambient conditions), U_g is the overall heat transfer 382 coefficient between the glass cover and the surroundings 383 (in this work, $U_g = 9 \text{ W/m}^2 \text{ K}$ chosen from the range 384 $1-15 \text{ W/m}^2 \text{ K}$ proposed by Garg (1987), h_{bw} is the convec-385 tive heat transfer coefficient between the black wall and the 386 air in the chimney and h_g is the convective heat transfer 387 coefficient between the glass cover and the air in the chim-388 ney. In addition, $a_{\rm bw}$ is the absorptance of the black wall 389 (in this work, $a_{\rm bw} = 0.9$, a value chosen from Duffie and 390 Beckman (1991), ε is the emmitance of the black wall (in 391 this work, $\varepsilon = 0.95$, a value chosen from Duffie and Beck-392 man (1991). Also, σ is the Stefan-Boltzmann constant 393 $(= 5.6697 \times 10^{-8} \text{ W/m}^2 \text{ K}^4)$ whereas $c_{p,\text{air}}$ and ρ_{air} are the 394 temperature dependent specific heat and density of air, 395 respectively. Finally, v is the average air velocity along 396 the chimney which since it can not stand as a fourth 397 unknown in the above system of equations it must be 398 described by some relation based on the other parameters 399 of the system (see below). 400

The convective heat transfer coefficients for both the 401 glass cover and the black wall and strictly for the vertical 402 position of the chimney are given by the relation (VDI-Wärmeatlas, 1991): 404

$$\begin{aligned}
\mathcal{M}_{u_{g,bw}} &= \frac{h_{g,bw}L}{\lambda} = \left\{ 0.825 + 0.387 \cdot \left(0.345 \cdot Ra_{g,bw} \right)^{1/6} \right\}^2 \\
\text{for } 10^{-1} < Ra\sin(s) < 10^{12} \end{aligned} \tag{38} 406$$



Fig. 2. Air velocity profile (a) and air temperature profile (b) across the chimney gap for different chimney lengths at the vertical position as calculated by the CFD model (day = 196, $H = 23.1 \text{ MJ/m}^2$, $T_{\text{amb}} = 28.9 \text{ °C}$).

а

Air velocity, m/s

b

ပ

temperature,

Air 30

20

0.000

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0.055

0.110

chimney gap for different tilt angles for a chimney of 1m length as calculated by the CFD model (day = 196, H = 23.1 MJ/m², $T_{\rm amb} = 28.9 \ ^{\circ}{\rm C}$).



Fig. 4. Air velocity profile (a) and air temperature profile (b), across the chimney gap for different tilt angles for a chimney of 4 m length as the CFD model (day = 196, $H = 23.1 \text{ MJ/m}^2$, calculated by $T_{\rm amb} = 28.9 \, {\rm °C}$).

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In the above, Nu is the Nusselt number, Ra is the Ray-407 leigh number, L is the length of the chimney and λ is the 408 thermal conductivity of air. For inclinations between 30° 409 and 75°, the heat transfer coefficient for the glass cover 410 h_{σ} (heated surface facing downwards) and the black wall 411 $h_{\rm bw}$ (heated surface facing upwards) are calculated from 412 the relations (VDI-Wärmeatlas, 1991): 413

$$Nu_{\rm g} = \frac{h_{\rm g}L}{\lambda} = 0.56[Ra_{\rm g}\sin(s)]^{1/4} \quad \text{for } 10^5 < Ra\sin(s) < 10^{11}$$
(39)

$$Nu_{bw} = \frac{h_{bw}L}{\lambda} = 0.56[Ra_{c}\sin(s)]^{1/4} + 0.13[Ra^{1/3} - Ra_{c}^{1/3}]$$

for $10^{8} < Ra\sin(s) < 10^{11}$ (40) 416

where Ra_c is a critical Rayleigh number that designates the 417 transition between laminar and turbulent flow and which is 418 419 given approximately by

$$\log(Ra_{\rm c}) = 8.9 - 0.00178 \cdot (90 - s)^{1.82} \tag{41}$$

Eqs. (39) and (40) were originally obtained from exper-422 iments with inclinations below 75°. For this, for inclina-423 tions between 75° and 90°, cubic spline interpolation is 424 employed to achieve smooth variation of coefficients with 425 inclination. 426

In order to describe the average air velocity inside the 427 chimney as a function of other system parameters, two dif-428 ferent expressions have been tried. The first one is derived 429



Fig. 5. Temperature of glazing (a) and absorber wall (b) along the normalized chimney length for different chimney lengths at the vertical position as calculated with the CFD model (day = 196, $H = 23.1 \text{ MJ/m}^2$, $T_{\rm amb} = 28.9 \ ^{\circ}{\rm C}$).

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430 by assuming that the pressure head inside a tilted chimney
431 counterbalances completely the pressure drop due to the
432 wall friction and inlet and outlet pressure losses. For equal
433 cross sectional areas at the inlet and outlet of the chimney
434 and for small density differences along the chimney this
435 yields:

$$f \cdot \frac{L}{D_{\rm H}} \cdot \frac{\rho_{\rm air}v^2}{2} + k_{\rm in} \cdot \frac{\rho_{\rm air}v^2}{2} + k_{\rm out} \cdot \frac{\rho_{\rm air}v^2}{2}$$

$$H_{\rm ch} \cdot g \cdot \sin(s) \cdot (\rho_0 - \rho_{\rm air}) \qquad (42)$$

438 where $D_{\rm H}$ is the hydraulic diameter of the chimney defined 439 as

$$441 \quad D_{\rm H} = \frac{2 \cdot w \cdot d}{w + d} \tag{43}$$

442 *w* is the width of the chimney gap and *d* is the depth of the 443 chimney gap. Also, k_{in} and k_{out} are the inlet and outlet 444 pressure loss coefficients, H_{ch} is the height difference be-445 tween outlet and inlet of the chimney (= $L \cdot sin(s)$) and *f* 446 is the wall friction coefficient calculated (for turbulent flow) 447 as

$$449 \quad f = \frac{0.316}{Re^{1/4}} \tag{44}$$

450 where *Re* is the apparent Reynolds number, defined as $D_{\rm H}$ 451 $v\rho_{\rm air}/\mu_{\rm air}$. Combining the above yields:

$$v = \left[\frac{2 \cdot L \cdot g(\sin(s))^2 (\rho_0 - \rho_{\rm air})}{\left(f \cdot \frac{L}{D_{\rm H}} + k_{\rm in} + k_{\rm out}\right) \cdot \rho_{\rm air}}\right]^{1/2}$$
(45)

455 For a rectangular channel with both ends open and 456 heated on one wall, Sandberg and Moshfegh (1998), pro-457 posed $k_{in} = 1.5$, $k_{out} = 1.0$ and f = 0.056.

The second expression that has been tried in the model 458 459 was described by Bansal et al. (1993) and Andersen 460 (1995). This is an empirical relation which uses the concept 461 of a discharge coefficient to adjust the air velocity for the 462 total flow resistances in the system (friction losses along the chimney wall, inlet and outlet pressure losses, etc). 463 464 For a case of equal cross sectional areas at the inlet and outlet of the chimney this relation reduces to (T in Kelvin): 465

468
$$v_{\text{ave}} = C_{\text{d}} \cdot \frac{\rho(T_{\text{air}})}{\rho(T_0)} \cdot \left[\frac{L \cdot g \cdot (\sin(s))^2 \cdot (T_{\text{air}} - T_0)}{T_0} \right]^{1/2}$$
(46)

469 where C_d is the discharge coefficient which for thermal 470 buoyant flows was proposed as 0.57 (Andersen, 1995).

471 2.2. CFD model

The commercial CFD code Fluent 6.1.18 is employed to simulate and check the heat transfer and fluid mechanics parts of the engineering model. For this, the CFD model uses as input data – apart from the chimney dimensions and material properties – the output data of the first two subroutines of the model, that is, the values of the total irradiation absorbed by the black wall and the glass cover.

479 Given the narrow geometry of our chimney (gap-tolength ratio 1:10), a 2D CFD model is considered adequate 480 based on the assumption of uniform temperature distribu-481 tions across the chimney width. The employed geometrical 482 domain has a variable length (1-12 m) as the first dimen-483 sion and a fixed gap depth (0.11 m) as the second one. 484 The third dimension (width = 0.74 m) is used only for esti-485 mation of total flow rates. The computational grid is a pure 486 map mesh with the cells clustered towards the black wall 487 and the glass. The grid for the 1 m high chimney consists 488 of 500 cells along the chimney and 55 cells across the gap 489 (27,500 guad cells in total), with an average size of 2 mm. 490 For the taller chimneys the grid size is increased propor-491 tionally in the length dimension to maintain the same spa-492 tial resolution. 493

Preliminary simulations showed that there are condi-494 495 tions where transition from laminar to turbulent flow occurs within the chimney and therefore, simulations are 496 performed with both the laminar and turbulent models. 497 For the latter, the shear-stress transport (SST) $k-\omega$ model 498 with the transitional flows option active is used (Fluent 499 user's guide, 2003), which is suitable for low Reynolds tur-500 bulent flows. This model combines the traditional two-501 layer turbulent zonal model with enhanced wall functions. 502 A fine mesh close to the walls is created with $y + \approx 2$, to 503 504 completely resolve the viscosity affected near-wall region.



Fig. 6. Temperature of glazing (a) and absorber wall (b) along the normalized chimney length for different chimney tilt angles for a chimney of 1 m length as calculated with the CFD model (day = 196, H = 23.1 MJ/m², $T_{amb} = 28.9$ °C).



Fig. 7. Average air velocity (a), average air temperature (b), average glazing temperature (c) and average absorber temperature (d) versus chimney length at the vertical position as calculated by two versions of the engineering model and the CFD code (day = 196, $H = 23.1 \text{ MJ/m}^2$, $T_{amb} = 28.9 \text{ °C}$).

505 The energy equation is employed to model the heat 506 transfer phenomena with the Boussinesg approximation 507 to hold for the density of air. Irradiation modeling is imple-508 mented using the Surface-to-Surface model (Fluent user's 509 guide, 2003), which accounts for the irradiation exchange 510 in an enclosure of gray-diffuse surfaces. The imposed 511 boundary conditions for the two chimney walls, (glazing 512 and absorbing black wall) are that they both have zero slip 513 and internal emmitance of 0.95.

514 3. Experiment

515 The experimental chimney duct has the shape of a nar-516 row parallelepiped with dimensions: 1 m height, 0.74 m 517 width and 0.11 m gap. Black painted aluminum sheet 518 (1.5 mm thick) is used for the construction of the rear 519 and side walls of the chimney. These walls have high solar 520 absorptance (~ 0.95) and low long wave emittance (~ 0.05) 521 (Garg, 1987). A 5 cm thick fiberglass layer ($\lambda = 0.045$ W/ 522 m K) is the outside insulation material of these walls. The 523 chimney's front side (glazing) is a commercial glass, 524 3 mm thick. The chimney glazing has a south orientation 525 at all times.

526 In order the chimney to stand at various inclinations, it 527 is mantled onto a short (0.5 m) metallic trapezoid equipped 528 with special fittings that allow the chimney to lie at different 529 tilt positions. Special care is given to make the chimney light enough so that it can be stably supported by the trapezoid when tilted and so permit easy handling of measuring probes. Even so, for slopes less than 45° it was difficult to hold the chimney firmly. 533

Along the two vertical narrow side walls of the parallel-534 epiped's section $(1 \times 0.11 \text{ m})$, five holes are drilled to facil-535 itate the insertion of measuring probes at distances 0.14 m, 536 0.34 m, 0.54 m, 0.75 m and 0.89 m from the bottom of the 537 chimney, respectively. Special contact-type thermocouples 538 (K type, OMEGA Inc.) are employed to measure the sur-539 face temperature of the glazing and the black wall at three 540 positions across the chimney width (left, center, right) to 541 check for 3D effects. From the recorded data, average val-542 ues are presented from those three positions since the var-543 iance (= SD/average) is less than 0.05. Fig. 1b shows a 544 photo of the constructed solar chimney where the five mea-545 suring stations along the vertical side wall are indicated. 546

Efforts have been made to measure the velocity of the air 547 in the chimney with a hot-wire anemometer probe fur-548 nished with sensitive temperature sensors (DO 2003, 549 550 DELTA OHM). Unfortunately, due to the low height of the chimney, the measured velocities were always below 551 552 0.2 m/s. For such low velocities – although within the measuring range of the anemometer – the readings were very 553 unstable perhaps due to minor atmospheric disturbances. 554 So, comparisons with theoretical predictions are based on 555 temperature measurements. This shortcoming is partly alle-556





Fig. 8. Average air velocity (a), average air temperature (b), average glazing temperature (c) and average absorber temperature (d) versus chimney tilt for a chimney of 1 m length as calculated by two versions of the engineering model and the CFD code (day = 196, $H = 23.1 \text{ MJ/m}^2$, $T_{amb} = 28.9 \text{ °C}$).

557 viated by the fact that with such a low height (i.e., light-558 weight) chimney it was possible to use it at different inclina-

559 tions, an issue essential for this work.

560 Total horizontal irradiation data are collected and inte-561 grated over 10 min intervals, for the period of the experi-562 ments with an Eppley Precision Pyranometer (model 563 PSP). The experiments are performed in Serres, Greece 564 (latitude 41°07', longitude 23°34', altitude 32 m).

565 4. Results and discussion

566 4.1. CFD parametric study

567 It is illustrative to display first the CFD calculations in order to appreciate the velocity and temperature profiles 568 569 in the chimney and their variation with respect to height 570 and tilt. Due to space limitations only simulations at a 571 summer day are presented: day 196 (mid July), monthly 572 average daily total irradiation on a horizontal plane 573 23.1 MJ/m^2 and monthly average daily ambient tempera-574 ture 28.9 °C. These data are taken for the city of Serres 575 from ELOT (1991).

576 Fig. 2a shows the air velocity profile across the chimney 577 gap at the exit of a vertical chimney, with chimney length 578 as a parameter. The shape of the velocity profiles for the 579 two smaller lengths (1 and 2 m) are typical of non-interacting boundary layers flowing past the absorber wall (gap 580 position = 0) and the glazing (gap position = 0.11), respec-581 tively. Two local maxima are observed near these walls (the 582 higher for the hotter absorber wall) whereas at the centre of 583 the chimney the velocity is close to zero. So, it is not so 584 strange that we were not able to measure significant air 585 velocities in our 1 m chimney. The situation changes for 586 higher chimneys where the two boundary layers start to 587 interact leading to less pronounced local maxima and a 588 smoother velocity front with appreciable velocities at the 589 center of the chimney. The overall air velocity (and there-590 591 fore air flow rate) increases significantly with chimney length due to the higher pressure head but also higher dif-592 593 ference between inside air temperature and ambient temperature. Above 4 m, full pipe flow prevails which, yet, is 594 not symmetrical across the gap. Inspection of the CFD 595 results (not shown due to space limitations) shows a tran-596 sition from laminar to turbulent flow for higher than 597 \sim 3 m chimneys. 598

Fig. 2b displays the corresponding mass-weighted – 599 "cup-mixing" – air temperature profiles. Mass-weighted 600 temperature values depict better the energy content of air 601 which affects the flow rate. As expected, the higher air temperatures are near the black absorber wall which seems to 603 be the main heat supplier of the system. Again, for 1 m and 604 2 m chimneys the two boundary layers hardly sense each 605



Fig. 9. Average air velocity (a), average air temperature (b), average glazing temperature (c) and average absorber temperature (d) versus chimney tilt for a chimney of 4 m length as calculated by two versions of the engineering model and the CFD code (day = 196, H = 23.1 MJ/m², $T_{amb} = 28.9$ °C).

606 other; the temperatures at the center of the gap being very 607 close to the incoming ambient temperature. This, changes 608 drastically for chimneys above 4 m. For the latter, an inter-609 esting change of the slope of the profiles occurs near the 610 glazing as a result of temperature mass-weighting.

611 Figs. 3 and 4 show the influence of the tilt position on 612 (a) air velocity and (b) mass-weighted air temperature 613 across the chimney gap for chimney lengths 1 m and 4 m, respectively. The main features of both the velocity and 614 temperature profiles are essentially those described in 615 Fig. 2 and do not seem to vary with tilt. For both chim-616 617 neys, it is clear that the higher velocities are achieved at 618 60° whereas the higher temperatures at 30°. In addition, 619 the air temperature in contact with the walls for the vertical 620 chimney is appreciably lower than the values for the other angles but this is not so for the velocity. Both the above 621 622 findings manifest a different influence of tilt on heat trans-623 fer and fluid flow in the chimney with the consequence that 624 the maximum energy uptake not to coincide with the max-625 imum air flow rate.

The influence of chimney length on (a) the glazing temperature and (b) the absorber temperature is shown in Fig. 5 for an inclination of 90°. Normalization in the length scale is performed by division with the total chimney length. Qualitatively speaking, the two walls exhibit similar trends. However, the absorber is always warmer than the glazing at the same length. Both walls are heated up signif-632 icantly within a very short distance from the inlet of the 633 chimney since there velocities are low and so energy cannot 634 be promptly transferred to the air stream. A bit upstream 635 where the walls are already warm, their temperature 636 increases more gradually because both the air velocity 637 and the temperature difference $(T_{wall} - T_{air})$ driving the 638 heat transfer towards the air become significant. Near the 639 top of the chimney, irradiation losses come into play and 640 reduce the temperature of the walls. It is interesting that 641 for chimneys between 1 and 4 m length, the local tempera-642 ture of the walls increases with length. On the contrary, for 643 higher chimneys the local temperature of the walls 644 decreases with length. This may be attributed to full pipe 645 turbulent flow that starts to develop for chimneys above 646 \sim 3 m. When this happens, the heat transfer coefficient 647 increases and so the walls are cooled down. The same tran-648 sition phenomena most probably explain also the stepwise 649 drop of temperature midway the 8 m chimney. For even 650 higher chimneys, e.g. 12 m, the wall temperatures increase 651 almost linearly with height indicating a pretty constant tur-652 bulent field. 653

Fig. 6 illustrates the influence of the tilt angle on (a) the 654 glazing temperature and (b) the absorber temperature for a 655 chimney of 1 m length. As can be seen, the local tempera-656 tures get higher as the tilt gets lower, for both walls. This is 657

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qualitatively what has been also observed in Fig. 3b regarding the air temperature and demonstrates that there is a direct relationship between the thermal condition of the walls and that of air. If one further considers Fig. 3a, it is apparent that in solar chimneys the tilt for maximum absorbed irradiation does not coincide with the tilt for maximum air flow.

665 4.2. Comparison between CFD and engineering models

666 Fig. 7a-d compares CFD predictions with predictions 667 from the engineering model, as a function of chimney length for a vertical orientation of the chimney. Predictions 668 669 refer to average air velocity and air temperature in the chimney as well as average glazing temperature and absor-670 671 ber temperature. Two series of model predictions are pre-672 sented; one based on Eq. (45) and the other on Eq. (46)673 for the estimation of air velocity in the chimney. Results 674 only for day 196 (mid July) are presented since this proved 675 to be the most stringent period for comparisons with the 676 largest deviations between models. In all four plots, it is 677 apparent that the two versions of the engineering model 678 give comparable results. Yet, they are different from the 679 CFD results. Regarding air velocity, CFD data are lower 680 than the engineering model data for chimneys less than 681 2 m but the situation reverses for chimneys above 4 m. 682 Air temperatures predicted by the engineering model are 683 below the values predicted by the CFD code for all the 684 examined lengths. Interestingly, CFD results show a non-685 monotonous sigmoid behavior with a kink point around 4 m. This is most likely due to the prevailing turbulent con-686 687 ditions for chimneys longer than 4 m. A non-monotonous 688 behavior is also observed in the CFD predictions of the 689 glazing and absorber temperatures with a peak value again close to 4 m. The latter means that for chimneys taller than 690 691 4 meters heat transfer from the walls to the flowing air is 692 drastically enhanced, indicating once more turbulent flow 693 conditions.

694 Figs. 8a-d and 9a-d show the dependence of all model predictions on the angle of the tilt for chimney lengths 695 1 m and 4 m, respectively. Calculations are again for day 696 697 198 (mid-July) where comparisons among models are less favorable. Despite the deviations among models in the pre-698 699 dicted values of air velocity, air temperature, glazing tem-700 perature and absorber temperature, there is a good 701 agreement on the optimum tilt that yields maximum air 702 velocity: The engineering model predicts an optimum tilt 703 around 65° whereas the CFD code around 60°. This result 704 has great significance as it lends support in the use of the 705 simpler engineering model for preliminary design purposes 706 and for comparisons between cases.

707 4.3. Comparison between predictions and experiments

Next, the CFD and the engineering model predictionsare compared against experimental measurements. Thedays for conducting the experiments were carefully selected

for wind speed to be less than 0.5 m/s. Runs were per-711 formed at days 305, 306 and 307 (beginning of November). 712 The values of total horizontal irradiation and ambient tem-713 perature mentioned in the Figure captions are those mea-714 sured on the spot and used as inputs to the models. It 715 must be stressed in advance that for November the devia-716 tions between CFD and engineering model predictions 717 are much less than for July (worst case). 718

Fig. 10a compares the predicted glazing and absorber 719 temperatures to measured values for a vertical position of 720 the chimney. Error bars denote the standard deviation of 721 measurements. During the measuring period the ambient 722 temperature was not constant so two series of CFD data 723 were calculated based on two different values for the ambi-724 ent temperature. The first one is the average temperature 725 and the second one the median temperature of the measur-726 ing period. As can be seen, the experimental data agree rea-727 sonably well with predictions. Fig. 10b displays 728 comparisons regarding the average air temperature. There 729 is again a fair agreement between data and predictions. 730 This is even more so if one considers that measurements 731 may be a bit higher than in reality due to the irradiation 732 absorbed by the finite size measuring probe (part of the 733



Fig. 10. Comparison between the engineering model predictions against CFD and experimental results as regards (a) the temperatures of the glazing and the absorber and (b) the air temperature along the chimney length. Plot (a): day = 307, $H = 8.12 \text{ MJ/m}^2$, median $T_{\text{amb}} = 19.2 \text{ °C}$, average $T_{\text{amb}} = 21.2 \text{ °C}$. Plot (b): day = 307, $H = 8.12 \text{ MJ/m}^2$, average $T_{\text{amb}} = 21.2 \text{ °C}$.

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Fig. 11. Comparison between the engineering model predictions against CFD and experimental results as regards the temperatures of the glazing and the absorber along the chimney length (a) for 90°, (b) 60° and (c) 45° angle of tilt. Plot (a): day = 305, $H = 9.70 \text{ MJ/m}^2$, $T_{amb} = 21.7 \text{ °C}$. Plot (b): day = 305, $H = 9.05 \text{ MJ/m}^2$, $T_{amb} = 19.1 \text{ °C}$. Plot (c): day = 306, $H = 9.90 \text{ MJ/m}^2$, $T_{amb} = 23.1 \text{ °C}$.

hot wire probe). Fig. 11a-c shows comparisons for the
chimney placed on the trapezoid base and fixed at three different inclinations. Again the agreement between predictions and measurements is good.

738 4.4. Chimney tilt for maximum air flow

Calculations with the engineering model (using Eq. (45))for the different months of the year to identify the optimum

tilt yielding maximum air flow are presented next, Fig. 12a. 741 The input values of monthly average daily total irradiation 742 on a horizontal plane and monthly average daily ambient 743 temperature, are taken from ELOT (1991). The corre-744 sponding maximum velocity values are also displayed to 745 allow appraisal of changes throughout the year. For com-746 parison, Fig. 12b shows the tilt that yields maximum 747 absorbed irradiation along with the corresponding irradia-748 tion values on the horizontal plane. 749

750 Clearly, air velocity and total irradiation exhibit similar trends receiving their lower values during summer months. 751 However, the angles themselves are very different. So, for 752 maximum air flow the chimney tilt varies in a rather nar-753 row range between 65° and 76° whereas for maximum irra-754 diation it varies between 12° and 44°. Furthermore, the 755 variation of irradiation values throughout the year is much 756 larger (on a percentage basis) than the variation of velocity 757 values. In the relevant work of Prasad and Chandra (1990), 758 conducted for a location in India, qualitatively similar 759 760 trends were observed but the range of the angles was different: between 53° and 76° for maximum air flow and 761 between 0° and 55° for maximum irradiation. 762

The question now arises on what is the best choice for 763 the tilt if the chimney is to be fixed at one and only inclina-764



Fig. 12. (a) Optimum tilt and corresponding maximum air velocity (for a 1 m and a 4 m chimney) and (b) optimum tilt and corresponding maximum total irradiation (for any chimney length) versus month of the year as calculated by the engineering model.

Please cite this article in press as: Sakonidou, E.P. et al., Modeling of the optimum tilt of a solar chimney for maximum air flow, Sol. Energy (2007), doi:10.1016/j.solener.2007.03.001

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Fig. 13. Average air velocity versus chimney tilt for a 1 m and a 4 m chimney as calculated by the engineering model for the mid day of December and July.

tion throughout the year. To help answer this, Fig. 13 dis-765 766 plays the air velocity versus the tilt for a 1 m and a 4 m 767 chimney and for the midday of December and July. Evi-768 dently, for winter applications the slight increase ($\sim 1\%$) 769 in air velocity by using the optimum tilt (as compared to 770 a vertical chimney) is not worth it against concerns about 771 the stability of the construction. However, during summer 772 months the gain in air velocity for the optimum tilt is 773 around 10% and decisions must be made more carefully.

774 5. Conclusion

775 A composite engineering model is developed that esti-776 mates the tilt of a solar chimney that yields the largest nat-777 ural air flow through it. The model starts by calculating the 778 solar energy absorbed by the solar chimney of varying tilt 779 and height for a given time (day of the year, hour) and 780 place (latitude). The monthly average daily value of total 781 irradiance and the ambient temperature are required as 782 inputs to the code along with some information on the 783 dimensions and properties of the construction materials (absorber, glazing, insulation). The outputs of the model 784 are the velocity and temperature of the air inside the chim-785 786 ney and the temperatures of the glazing and the black painted absorber, as a function of tilt and height. Compar-787 788 isons of the model predictions with CFD results for a 789 broad range of chimney lengths (1-12 m) and tilts 790 $(30-90^{\circ})$ delineates the usefulness of the model but marks 791 also its limitations. Moreover, model predictions are in 792 satisfactory accord with experimental measurements from 793 a 1 m chimney operated at different inclinations. The rea-794 sonable agreement between model predictions with CFD 795 and experimental results encourages the use of the engi-796 neering model as a tool for evaluating design parameters 797 and for comparative studies.

798 6. Uncited reference

799 Ding et al. (2005).

Acknowledgements

The project is co-funded by the European Social Fund 801 and National Resources – (EPEAEK-II) ARHIMIDIS. 802 The authors are indebted to Mr. N. Vallous for his editorial work. 804

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